

Spectral principal component analysis of mid-infrared spectra of a sample of PG QSOs

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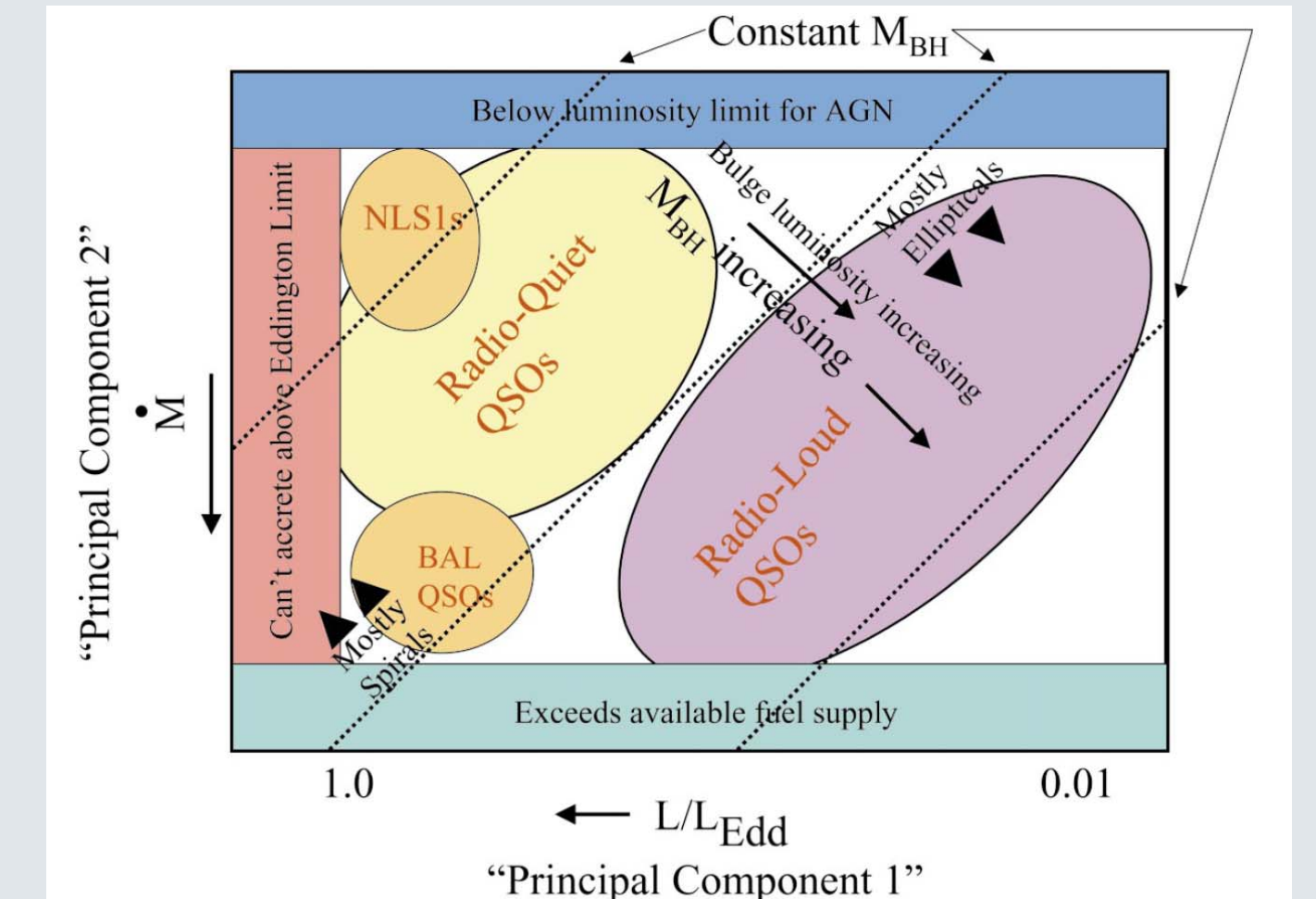
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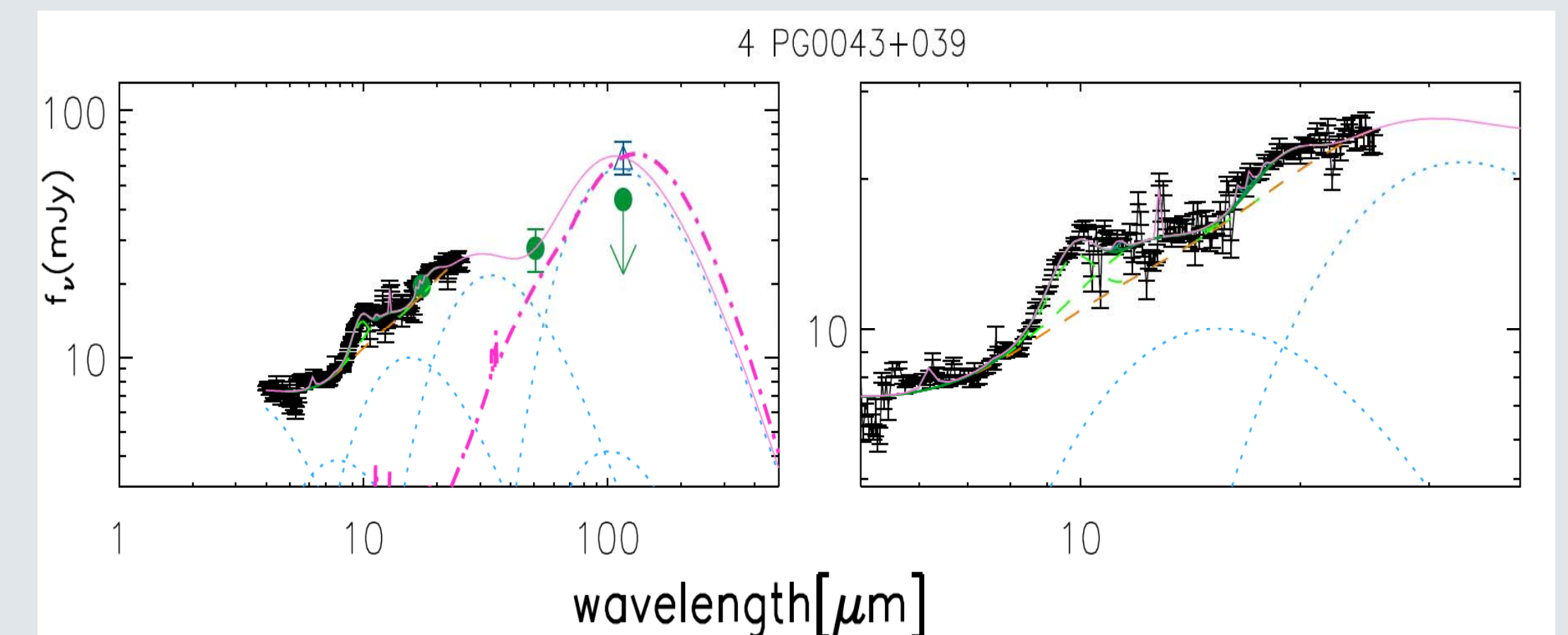
1. Introduction

- AGN/QSOs: standard paradigm, SMBH, BLRs/NLRs, torus, jet. Radiation + orientation => SED diversity.
- Torus+host: Si, PAH, molecular H, ion; SF, 11.3 μm PAH, MIPS 70 μm , MIPS/PACS 160 μm
- PCA: multivariate correlation analysis; find the fewest number of variables (PCs) to describe the variance. Drivers of PC1/PC2: PC1-> L/Ledd, PC2 -> L



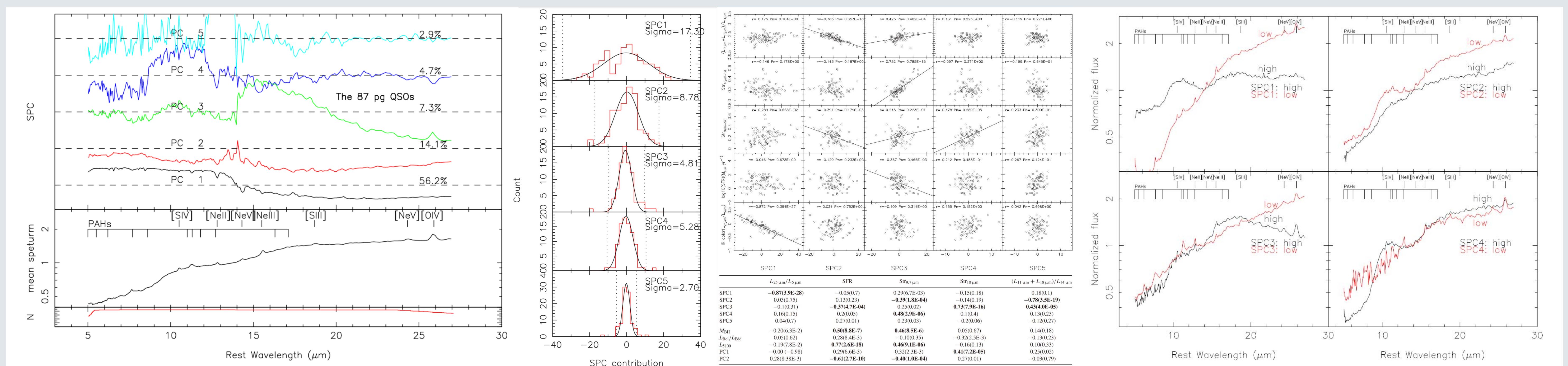
2. Sample

- 87 PG QSOs: $z < 0.5$; mid-infrared, Spitzer IRS 5-40 μm , MIPS 24, 70, 160 μm , Herschel PACS 160 μm ; SMBH mass, Eddington ratio
- Model IRS spectra+ MIPS/PACS: torus+SF, (multi-T BB, silicate, aromatic, emission lines, the SF template), 26K-1000K (11 BB), two two-Gaussian fit for Si 9.7, 18, Drude profiles for PAHs, one-Gaussian fit for other emission lines.

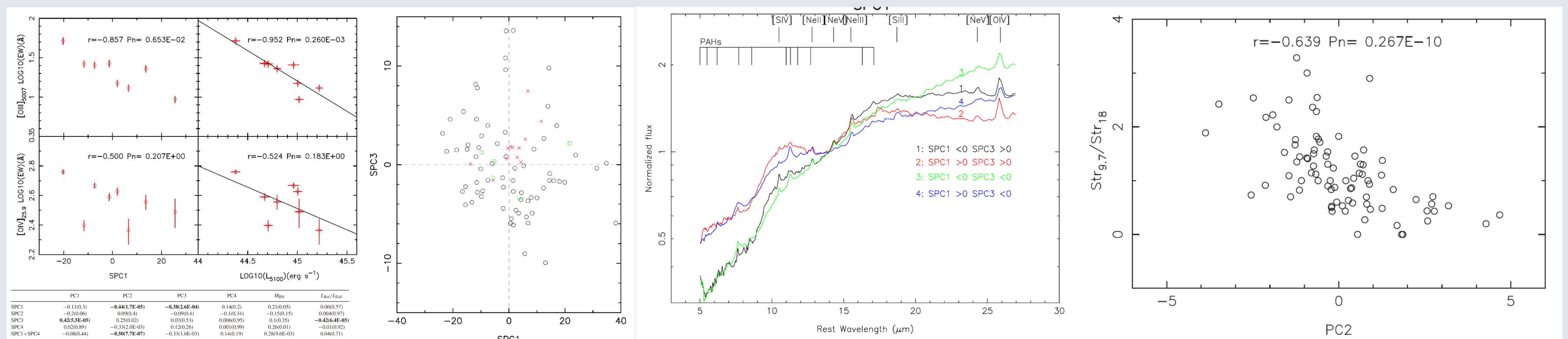


3. Results

- The first five mid-IR eigenspectra: 85.2%, relations (SPC1/2/3/4/5 versus IR-slope, SFR, Si); mean spectra for $|SPC_{ij}| > 1\sigma$



- The relation with PCs by BG92: PC1/2/3/4 + M_{BH} + L/L_{Edd} , SPC1-SPC3



- The connection between SFR and PCs: more Luminosity -> larger SFR, feedback by MBH accretion.

4. Summary

- SPCA of mid-IR spectra of PG QSOs, first five eigenspectra, 85.2%.
- SPC1: mid-IR slope, forbidden IR lines EW, Baldwin effect, Larger AGN luminosity -> more dust, smaller opening angle -> suppressed EW
- SPC3: Si 18 μm , SFR(PAH 11.3 μm); SPC4: Si 9.7 μm
- SPCs-PCs: SPC1-PC2, SPC3-PC1; Strong correlation between Si9.7/Si18 and PC2 (L_{5100} , MBH, accretion rate)
- SFR-PCs: correlation between SFR and PC2/MBH, possible feedback by SMBH accretion.

Reference

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- Shi Y., Rieke G. H., Ogle P. M., Su K. Y. L., Balog Z., 2014, ApJS, 214, 23